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Recent advances in the theory of finite correlated systems: From few-body scattering to phase transitions

J. Berakdar

Max-Planck Institut für Mikrostrukturphysik, Weinberg 2, 06120 Halle, Germany

ABSTRACT

This work gives a brief overview on the recent progress in the analytical treatment of correlated fewbody Coulomb continuum systems. Special emphasis is put on the approximate separability of the long and short-range dynamics. This separability is exposed by expressing the total Hamiltonian of the system in appropriate curvilinear coordinates in which case the total Hamiltonian breaks down into two commuting operators: An operator whose eigenstates decay for large inter-particle distances. The states associated with the second operator possess an oscillatory behaviour in the asymptotic region and can thus be assigned to the long-range behaviour of the few-body system.

At finite distances the total Hamiltonian contains in addition to these two operators a term which mixes the long and the short range dynamics.

The many-body wave function in the asymptotic region is derived and discussed. Methods that couple the asymptotic region to finite distances are also presented. The strength and weaknesses of the derived approximate wave functions are demonstrated by evaluating scattering transition matrix elements and comparing to available experimental data. We also discuss a recent Green



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function methodology that is capable of dealing with large finite and extended systems of charged particles. In addition, we give a brief account of the Green function approach as applied for the determination of the thermodynamics properties and critical phenomena of finite interacting systems.

2. INTRODUCTION

There are basically two fundamental difficulties in the theoretical treatment of systems consisting of few charged particles. The first one stems from the inherent non-separability of many-body interacting systems. This problem appears basically for all forms of the inter-particle interactions (apart from special cases, such as contact two-body potentials). In addition, theories for charged particle systems have to deal with the infinite range of the Coulomb interaction that precludes free asymptotic states and hence limits seriously the applicability of standard many-body methods.

A prototypical example of many-particle correlated systems is the interacting electron gas (interacting jellium model). In their seminal treatment of this system Pines and coworkers [1] have pointed out that a dense, interacting electron gas can be described quantum mechanically by expressing the long-range part of the inter-electronic Coulomb interactions in terms of collective fields.

These fields represent organized plasma oscillations of the electron gas as a whole. The total Hamiltonian can then be written in a form that describes these collective modes and a set of individual electrons which interact with one another via short-range screened Coulomb potentials. The short-range part of the electron-electron interaction can be parameterized remarkably well by a Yukawa-type potential (exponentially screened Coulomb potential) with a screening length depending on the density of the electrons. There is, in addition, a mixing term that couples the individual particles to the collective modes. This mixing

term can be eliminated under certain conditions [1].

This insightful knowledge that the long-range dynamics can be decoupled from the short-range one may seem somewhat restricted to the high density limit where the mean kinetic energy dominates over the mean potential energy. In fact in the dilute gas case where the electronic correlation is the dominating factor new physical phenomena set in such as the particle localization known as the Wigner crystallization. Therefore it is not clear whether the concepts developed in the realm of condensed matter theory will hold true for few-particle systems.

In this work it will be shown that in a finite electronic system and under certain conditions specified below, it is indeed possible to decouple the long-range part of the Hamiltonian from the short range part. This is achieved by an analytical consideration of the eigenstates of the many-body, non-relativistic Schrödinger equation without going through the procedures and difficulties of the scattering theory for charged particles.

In recent years, much efforts have been focused towards the asymptotic separability of the Schrödinger equation in the fragmentation channel [2-13], i.e. the solution of the Schrödinger equation is sought for a large hyperradius (size) of the system. Occasionally the wave functions derived in this way are called asymptotic wave functions. It should be remarked however that while these wave functions are asymptotically correct, they are defined in the entire Hilbert space and therefore they can be employed for the evaluation of transition amplitudes. It is only that in the asymptotic region the coupling to short range dynamics vanishes and therefore the asymptotic treatment becomes justified. In this context it is important to note that the asymptotic region is reached for large interparticle separations or at high particles' energies. This means, for sufficiently high energies, the asymptotic regime extends to small inter-particle distances to cover a considerable range of the Hilbert space.

At short distances and low energies, the short-range part of the Hamiltonian and it's coupling to the asymptotic region might play a dominant role and a treatment, whose justification is based only on asymptotic arguments is not appropriate.

In this article we analyze the scattering states of a finite electronic systems starting from the two-body collision and closing the discussion by a brief review of the few-body Green function method and its application to thermodynamic properties and critical phenomena in finite systems.

Unless otherwise stated we employ atomic units throughout and neglect relativistic corrections.

3. TWO-BODY COULOMB SCATTERING

For a transparent treatment it is important to express the Schrödinger equation in an appropriate coordinate system. To sense the "natural" coordinates for fragmentation processes involving Coulomb potentials it is instructive to consider the non-relativistic scattering of two charge particles with charges z_1 and z_2 . The Schrödinger equation describing the motion in the two-particle relative coordinater $\bf r$ is

$$\left[-\frac{1}{2\mu} \Delta + \frac{z_1 z_2}{r} - E \right] \Psi_{\mathbf{k}}(\mathbf{r}) = 0. \tag{1}$$

Here **k** is the momentum conjugate to **r** and $E = k^2/2\mu$ is the energy whereas μ is the reduced mass of the two particles. To decouple kinematics from dynamics we make the anastz:

$$\Psi_{\mathbf{k}}(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}}\bar{\Psi}_{\mathbf{k}}(\mathbf{r}). \tag{2}$$

The effect of the potential is totally described by the term $\bar{\Psi}$ in Eq.(2). To inspect the asymptotic properties of (1) we substitute (2) in (1) and disregard terms that fall off faster than the Coulomb potential which yields the equation

$$\left[-\frac{i}{\mu} \mathbf{k} \cdot \nabla + \frac{z_1 z_2}{r} \right] \bar{\Psi}_{\mathbf{k}}(\mathbf{r}) = 0.$$
 (3)

This equation can be solved by the ansatz $\bar{\Psi} = \exp(i\phi)$ which, upon insertion in Eq.(3), leads to the Coulomb-phases

$$\phi_{\mathbf{k}}^{\pm}(\mathbf{r}) = \pm \frac{z_1 z_2 \mu}{k} \ln a(r \mp \hat{\mathbf{k}} \cdot \mathbf{r}).$$
 (4)

The factor $z_1z_2\mu/k$ is called the Sommer-feld parameter and is an indicator for the strength of the interaction. The integration constant a has a dimension of a reciprocal length and is usually set to be a=k. The key point for the present work is that the coordinate inherent to Coulomb scattering is the so-called *parabolic* coordinate $\xi^{\pm} := r \mp \hat{\mathbf{k}} \cdot \mathbf{r}$ where the + or - sign should be chosen if one is dealing with incoming or outgoing-wave boundary conditions.

4. THREE-BODY COULOMB SCATTERING

For three-body continuum systems the situation is much more complex since the Schrödinger equation is not separable. Nonetheless, as a first step, one might think of a three-body system as the subsume of three non-interacting two-body subsystems [9]. Since we know the appropriate coordinates for each of these two-body subsystems, as illustrated above, the obvious choice of coordinates would be

$$\{\xi_k^{\mp} = r_{ij} \pm \hat{\mathbf{k}}_{ij} \cdot \mathbf{r}_{ij}\},$$

$$\epsilon_{ijk} \neq 0; \ j > i, k \in [1, 3],$$
(5)

where \mathbf{r}_{ij} is the position of the particle i relative to the particle j and $\hat{\mathbf{k}}_{ij}$ denote the directions of the momenta \mathbf{k}_{ij} that are conjugate to r_{ij} . Since we are dealing with a six-dimensional problem three other independent coordinates are needed in addition to (5). To make a

reasonable choice for these remaining coordinates we remark that, usually, the momenta \mathbf{k}_{ij} are determined experimentally, i. e. they can be considered as the laboratory fixed coordinates. In fact it can be shown that the coordinates (5) are related to the Euler angles. Thus, it is advantageous to choose body-fixed coordinates. Those are conveniently chosen as

$$\{\xi_k = r_{ij}\}, \quad \epsilon_{ijk} \neq 0; \ j > i, k \in [4, 6].$$
 (6)

Upon a mathematical analysis it can be shown that the coordinates (5,6) are linearly independent [9] except for some singular points where the Jacobi determinant vanishes. The main task is now to rewrite the three-body Hamiltonian in the coordinates (5,6). After factoring out the trivial plane-wave part [as done in Eq.(2)] it turns out that the three body wave function is determined as an eigensolution of an operator H with zero eigen value- [9]. The differential operator H, expressed in the curvilinear coordinates (5,6), has the structure

$$H = H_{par} + H_{int} + H_{mix}. (7)$$

The operator H_{par} is differential in the parabolic coordinates $\xi_{1,2,3}$ only whereas H_{int} acts on the internal degrees of freedom $\xi_{4,5,6}$. The mixing term H_{mix} arises from the off-diagonal elements of the metric tensor and plays the role of a rotational coupling in a hyperspherical treatment.

The essential point is that the differential operators H_{par} and H_{int} are exactly separable in the coordinates $\xi_{1...3}$ and $\xi_{4...6}$, respectively, for they can be written as [9]

$$H_{par} = \sum_{j=1}^{3} H_{\xi_j}; \ [H_{\xi_j}, H_{\xi_i}] = 0;$$

$$\forall \ i, j \in \{1, 2, 3\},$$
(8)

and

$$H_{int} = \sum_{j=4}^{6} H_{\xi_j}; \ [H_{\xi_j}, H_{\xi_i}] = 0;$$

$$\forall i, j \in \{4, 5, 6\},\tag{9}$$

where

$$H_{\xi_{j}} = \frac{2}{\mu_{lm}r_{lm}} \left[\partial_{\xi_{j}} \xi_{j} \partial_{\xi_{j}} + ik_{lm} \xi_{j} \partial_{\xi_{j}} - \mu_{lm} z_{lm} \right];$$

$$\epsilon_{jlm} \neq 0, \ j \in \{1, 2, 3\}.$$
(10)

and

$$H_{\xi_4} = \frac{1}{\mu_{23}} \left[\frac{1}{\xi_4^2} \partial_{\xi_4} \, \xi_4^2 \, \partial_{\xi_4} + i2k_{23} \, \frac{\xi_1 - \xi_4}{\xi_4} \partial_{\xi_4} \right]; (11)$$

$$H_{\xi_5} = \frac{1}{\mu_{13}} \left[\frac{1}{\xi_5^2} \partial_{\xi_5} \, \xi_5^2 \, \partial_{\xi_5} + i2k_{13} \, \frac{\xi_2 - \xi_5}{\xi_5} \partial_{\xi_5} \right]; \quad (12)$$

$$H_{\xi_6} = \frac{1}{\mu_{12}} \left[\frac{1}{\xi_6^2} \partial_{\xi_6} \, \xi_6^2 \, \partial_{\xi_6} + i2k_{12} \, \frac{\xi_3 - \xi_6}{\xi_6} \partial_{\xi_6} \right]. \tag{13}$$

In the equations above the reduced mass of the pair ij is denoted by μ_{ij} and its product charge by z_{ij} . The operator $H_{mix} = H - H_{par}$ derives from the expression

$$\begin{array}{ll} H_{mix} := & \sum_{u \neq v=1}^{6} & \left\{ (\nabla_{\mathbf{r}_{ij}} \xi_{u}) \cdot (\nabla_{\mathbf{r}_{ij}} \xi_{v}) \right. \\ & \left. + (\nabla_{\mathbf{R}_{k}} \xi_{u}) \cdot (\nabla_{\mathbf{R}_{k}} \xi_{v}) \right\} \partial_{\xi_{u}} \partial_{\xi_{v}} \end{array} \tag{14}$$

where \mathbf{R}_k indicates the position of the center of mass of the pair ij with respect to the particle k.

Noting that H_{ξ_j} , j=1,2,3 is simply the Schrödinger operator for two-body scattering rewritten in parabolic coordinates (after factoring out the plane-wave part), one arrives immediately, as a consequence of Eq.(8), at an expression for the three-body wave function as a product of three two-body continuum waves, provided that H_{int} and H_{mix} are negligible. Fortunately, it turns out that the matrix elements of H_{int} and H_{mix} are in fact small compared to those of H_{par} in case of large interparticle separations [9] or at high particles' energies. It should be emphasized, however, that this "asymptotic separability" is not the result of using the coordinate system (5). It is

only that the operator H_{par} attains in the coordinates (5) a very simple and transparent form for an arbitrary three-body system. In fact, this same operator H_{par} has a much more complex representation in the usual Jacobi coordinates (cf. Ref. [9]).

With decomposing the total Hamiltonian in H_{par} , H_{int} and H_{mix} we achieved a result similar to that obtained by Pines and co-worker for the electron gas: The system is decomposed into a long-range and a short range components described respectively by H_{par} and H_{int} . As shown below, the eigen-function of H_{par} has an oscillatory asymptotic behaviour whereas the eigenstates of H_{int} decay for large interparticle distances [14]. The mixing term H_{mix} couples the short range to the long-range modes of the system.

A popular method to obtain three-body continuum wave functions has been to diagonalize H_{par} and neglect all other terms in the Hamiltonian H_{int} and H_{mix} . It is clear from the above discussion however, that when physical situations are considered in which the short-range dynamics is important it will be insufficient to treat only the long-range part of the three-body wave function, i.e. to approximate this wave function by the eigenfunctions of H_{par} . For example, In the extreme case where all particles are close to each other the terms H_{int} and H_{mix} are the important ingredient of H whereas H_{par} is then irrelevant. Hence, it is comprehensible that the behaviour of the eigenfunctions of H_{par} is at variance (cf. [9, 16]) with the Fock expansion [15] that has to be satisfied at the three-body collision point.

4.1. REMARKS ON THE APPROXIMATE SEPARABILITY

The analytical structure of the Eqs.(8-14) warrants several remarks:

The total potential is contained in the operator H_{par} , as can be seen from Eqs.(10).

Thus, the eigenstates of H_{par} treats the total potential in an exact manner. This means on the other hand that the operators H_{int} and H_{mix} are parts of the kinetic energy operator.

This situation is to be contrasted with other treatments [17-25] of the three-body problem in regions of the space space where the potential is smooth, e.g. near a saddle point. In this case one usually expands the potential around the f point and accounts for the kinetic energy in an exact manner.

In Eq.(10) the total potential appears as a sum of three two-body potentials. It should be stressed, that this splitting is arbitrary, since the dynamics is controlled by the total potential. I.e., any other splitting that leaves the total potential invariant is equally justified. This fact we will use below for the construction of three-body states. For large inter-particle separation the operators H_{int} and H_{mix} are negligible as compared to H_{par} and the splitting of the total potential as done in Eqs.(10) becomes unique. This means, for large particles' separation the three-body dynamics is controlled by sequential two-body scattering events.

The momentum vectors \mathbf{k}_{ii} enter the Schrödinger equation via the asymptotic boundary conditions. Thus, their physical meaning, as two-body relative momenta, is restricted to the asymptotic region of large inter-particle distances. The consequence of this conclusion is that, in general, any combination or functional form of the momenta \mathbf{k}_{ii} is legitimate as long as the total energy is conserved and the boundary conditions are fulfilled (the energies and the wave vectors are liked via a parabolic dispersion relation). This fact has been employed in Ref. [6] to constructed threebody wave functions with position-dependent momenta kij and in Ref. [26] to account for off-shell transitions.

The separability of the operators (11-13) may be used to deduce representations of three-body states [14] that diagonalize simultaneously H_{par} and H_{int} . It should be noted however, that generally the operator H_{mix} , which has to be neglected in this case, falls off with distance as fast as H_{int} .

As well-known, each separability of a system implies a related conserved quantity. In

the present case we can only speak of an approximate separability and hence of approximate conserved quantum numbers.

If we discard H_{int} and H_{mix} in favor of H_{para} , which is justified for $k_{ij} \, \xi_k, \, \in_{ijk} \neq 0, \, k \in [1, 3]$ (i.e. for a large ξ_k or for a high two particle momentum k_{ij}), the three-body good quantum numbers are related to those in a two-body system in parabolic coordinates. The latter are the two-body energy, the eigenvalue of the component of the Lenz-Runge operator along a quantization axis z and the eigenvalue of the component of the angular momentum operator along z. In our case the quantization axis z is given by the linear momentum direction $\hat{\mathbf{k}}_{ij}$.

In Ref. [9] the three-body problem has been formulated in hyperspherical-parabolic coordinates. In this case the operator H_{int} takes on the form of the grand angular momentum operator. This observation is useful to expose the relevant angular momentum quantum numbers in case H_{mix} can be neglected.

In Ref. [5, 27] the three-body system has been expressed in the coordinates $\eta_i =$ ξ_{j}^{+} , j = 1,2,3 and $\overline{\eta}_{j} = \xi_{j}^{-}$, j = 1,2,3. This is the direct extension of the parabolic coordinates for the body problem (cf. Section 3) to the three-body problem. From a physical point of view this choice is not quite suited, for scattering states are sufficiently quantified by outgoing or incoming wave boundary conditions (in contrast to standing waves, such as bound states whose representation requires a combination of incoming and outgoing waves). Therefore, to account for the boundary conditions in scattering problems, either the coordinates η_i or $\overline{\eta}_i$ are needed. The appropriate choice of the remaining three coordinates should be made on the basis of the form of the forces governing the three-body system. In the present case where external fields are absent we have chosen $\xi_k = r_{ij}$, k = 4, 5, 6 as the natural coordinates adopted to the potential energy operator.

4.2. COUPLING THE SHORT AND THE LONG-RANGE DYNAMICS

In the preceding sections we pointed out that the eigenstates of H_{par} can be deduced analytically. These eigenfunctions, even though are well defined in the entire Hilbert space, constitute a justifiable approximation to the exact three-body state in the asymptotic region only (e.g. in the region of larger inter-particle separation or for higher energies). This fact is important when it comes to evaluating reaction amplitudes, for such amplitudes involve the many-body scattering state in the entire Hilbert space. Therefore, an adequate description of the short-range dynamics may be necessary, in particular in cases where the contributions to the matrix elements of the transition amplitudes originates from the internal region, i.e. when the reaction takes place at small interparticle distances. Nevertheless the eigenstates of H_{par} can, and have been used for the calculations of transition matrix elements, as for example done below. In this case the justification of this doing must go beyond the asymptotic correctness argument.

In this section we seek three-body wave functions that diagonalize, in addition to H_{par} , parts of H_{int} and H_{mix} .

One method that turned out to be particularly effective for this purpose relies on the observations: a) In a three-body system the form of the two-body potentials z_{ij}/r_{ij} are generally irrelevant, as long as the total potential is conserved. c) To keep the mathematical structure of the operators (8,10) unchanged and to introduce a splitting of the total potential while maintaining the total potential's rotational invariance one can assume the strength of the individual two body interactions, characterized by z_{ij} , to be dependent on $\xi_{4,5,6}$. This means we introduce position dependent product charges as

$$\bar{z}_{ij} = \bar{z}_{ij}(\xi_4, \xi_5, \xi_6),$$
 (15)

with

$$\sum_{j>i=1}^{3} \frac{\bar{z}_{ij}}{r_{ij}} = \sum_{j>i=1}^{3} \frac{z_{ij}}{r_{ij}}.$$
 (16)

To obtain the *many-body* potentials $\overline{V}_{ij} := \overline{z}_{ij} / r_{ij}$ we express them as a linear mixing of the isolated *two-body* interactions $V_{ij} := z_{ij} / r_{ij}$, i.e

$$\begin{pmatrix}
\overline{V}_{23} \\
\overline{V}_{13} \\
\overline{V}_{12}
\end{pmatrix} = \mathcal{A} \begin{pmatrix} V_{23} \\
V_{13} \\
V_{12} \end{pmatrix},$$
(17)

where $\mathcal{A}(\xi_4, \xi_5, \xi_6)$ is a 3×3 matrix. The matrix elements are then determined according to 1) the properties of the total potential surface, 2) to reproduce the correct asymptotics of the three-body states and 3) in a way that minimizes H_{int} and H_{mix} . It should be stressed that the procedure until this stage is exact. It is merely a splitting of the total potential that leaves this potential and hence the three-body Schrödinger equation unchanged.

4.3. AN ELECTRON PAIR IN THE FELD OF A POSITIVE ION

To be specific let us demonstrate the method for the case of two electrons moving in the Coulomb field of a residual ions. This brings about some simplifications since the ion can be considered infinitely heavy as compared to the electron mass. Traditionally, the electrons are labeled a and b and their positions and momenta with respect to the residual ion are respectively called \mathbf{r}_a , \mathbf{r}_b and \mathbf{k}_a , \mathbf{k}_b . Adopting this notation, the eigenstate of the operator H_{par} reads:

$$\overline{\Psi}_{\mathbf{k}_{a},\mathbf{k}_{b}}(\xi_{1\cdots 6}) = {}_{1}F_{1}(i\beta_{a}, 1, -ik_{a} \xi_{1})
{}_{1}F_{1}(i\beta_{b}, 1, -ik_{b} \xi_{2})
{}_{1}F_{1}(i\beta_{ab}, 1, -ik_{ab} \xi_{3}).$$
(18)

Here we denoted the relative electron-electron momentum by $\mathbf{k}_{ab} = \frac{1}{2}(\mathbf{k}_a - \mathbf{k}_b)$ whereas ${}_{1}F_{1}[a,b,x]$ stands for the confluent

hypergeometric function and β_j , $j \in \{a, b, ab\}$ are the sommerfeld parameters

$$\beta_j = \frac{\bar{z}_j}{v_j}, \ j \in \{a, b, ab\},\tag{19}$$

with v_j being the velocities corresponding to the momenta \mathbf{k}_j and z_j , $j \in \{a, b, ab\}$ are the electrons-ion and electron-electron effective product charges, respectively. The form of \overline{z}_j , j is still to be determined.

Below the functions \bar{z}_j ($\xi_{4...6}$) are given that preserve the total potential, possess the correct three-body asymptotics and incorporate features of the many-particle motion at the complete fragmentation threshold, namely along the saddle point ofthe total potential, the so-called Wannier ridge [17-25]. Since \bar{z}_j are assumed to depend on the internal co-ordinates only (ξ_4 , ξ_5 , ξ_6) the wave function (18) is still an eigenstate of the long-range Hamiltonian H_{par} (given by Eq.10). In physical terms it can be said that the effect of the short-range part of the Hamiltonian is to modify dynamically the coupling strength of the isolated two particle system (z_{ij}).

To ensure the invariance of the Schrödinger equation under the introduction of the product charges $\bar{z}_j(\xi_{4...6})$ the three conditions

$$\sum_{j} \frac{\bar{z}_{j}(\xi_{4\cdots 6})}{r_{j}} \equiv \frac{-z}{r_{a}} + \frac{-z}{r_{b}} + \frac{1}{r_{ab}},$$

$$j \in \{a, b, ab\},$$
(20)

have to be fulfilled (\mathbf{r}_{ab} is the electron-electron relative coordinate and z is the charge of the residual ion). The wave functions containing \bar{z}_j must be compatible with the three-body asymptotic boundary conditions. These are specified by the shape and size of the triangle formed by the three particles (two electrons and the ion): I.e., the derived wave function must be, to a leading order, an asymptotic solution of the three-body Schrödinger

equation when the afrorementioned triangle tends to a line (two particles are close to each other and far away from the third particle) or in the case where, for an arbitrary shape, the size of this triangle becomes infinite. The latter limit implies that all interparticle co-ordinates $r_{a,b,ab}$ must grow with the same order, otherwise we eventually fall back to the limit of the three-particle triangle being reduced to a line [9], as described above. In addition we require the Wannier threshold law for double electron escape to be reproduced when the derived wave functions are used for the evaluation of the matrix elements.

All of the above conditions are sufficient to determine \bar{z}_j and thus the wave function (18). This wave function is called "dynamically screened three-body Coulomb wave function" $\Psi_{\rm DS3C}$. This is because this wave function consists formally of three Coulomb waves where the short-range dynamics enters as a dynamical screening of the strength of the two-body interaction.

The applicability of the wave function $\Psi_{\rm DS3C}$ to scattering reactions is hampered by the involved functional dependence leading in the numerical complications determination of the normalization and of the scattering matrix elements. Furthermore, the incorporation of the three-body scattering dynamics at shorter distances brings about intrinsic practical disadvantages as compared to an approach where \bar{z}_j are constant (this approach is usually called three-body Coulomb wave method, 3C). Namely, the construction of Ψ_{DS3C} has to be individually undertaken for given charge and mass states of the specific three particle system at hand. This is comprihensible since properties of potential are inherent to the the total threebody under particular system investigation.

As shown in Ref. [28] the normalization of the 3C wave function is readily determined from the asymptotic flux. This procedure has not been accessible in the case

of $\Psi_{\rm DS3C}$ due to the position dependence of \overline{z}_i .

To overcome this difficulty (and that associated with the six-dimensional numerical integration when evaluating transition matrix elements) we note that the position dependence of \bar{z}_j (r_a , r_b , r_{ab}) occurs (due to dimensionality considerations) through ratios of the interparticle distances. Thus, this dependence can be converted into velocity dependence by assuming that

$$\frac{r_i}{r_j} \propto \frac{v_i}{v_j}$$
. (21)

The proportionality constant in Eq. (21) could be of an arbitrary functional dependence. It should be emphasized that the approximation (21) is not a classical one, i. e. it is not assumed that the particles' motions proceed along classical trajectories [conversely, if the motion were classically free, Eq. (21) holds]. It merely means that the total potential is exactly diagonalized in the phase space where Eq. (21) is satisfied, as readily deduced from Eq. (20).

Eq. (21) renders possible the normalization of Ψ_{DS3C} since in this case we obtain $\bar{z}_j = \bar{z}_j$ (k_a , k_b , k_{ab}) and the arguments used in Ref. [28] can be repeated to deduce for the nor-malization N the expression

$$N = \prod_{j} N_{j}, j \in \{a, b, ba\}$$

$$N_{j} = \exp[-\beta_{j}(k_{a}, k_{b}, k_{ba})\pi/2]$$

$$\Gamma[1 - i\beta_{j}(k_{a}, k_{b}, k_{ba})]. \tag{22}$$

Here Γ (x) is the Gamma functions. The velocity dependent product charges [29] have the form

$$\bar{z}_{ba}(\mathbf{v}_a, \mathbf{v}_b) = [1 - (f \ g)^2 \ a^{b_1}] a^{b_2}$$
 (23)

$$\bar{z}_a(\mathbf{v}_a, \mathbf{v}_b) = -1 + (1 - \bar{z}_{ba}) \frac{v_a^{1+a}}{(v_a^a + v_b^a)v_{ab}}$$
 (24)

$$\bar{z}_b(\mathbf{v}_a, \mathbf{v}_b) = -1 + (1 - \bar{z}_{ba}) \frac{v_b^{1+a}}{(v_a^a + v_b^a)v_{ab}}.$$
 (25)

The functions occurring in Eqs. (23,24) are defined as $(\mathbf{v}_a, \mathbf{v}_b)$ are the electron's velocities and $\mathbf{v}_{ab} = \mathbf{v}_a - \mathbf{v}_b$

$$f := \frac{3 + \cos^2 4\alpha}{4}, \quad \tan \alpha = \frac{v_a}{v_b}$$
 (26)

$$g := \frac{v_{ab}}{v_a + v_b} \tag{27}$$

$$b_1 := \frac{2v_a v_b \cos(\theta_{ab}/2)}{v_a^2 + v_b^2} \tag{28}$$

$$b_2 := g^2(-0.5 + \bar{\mu})$$
 E
(29)

$$a := \frac{1}{E + 0.5}, \tag{30}$$

where E is being measured in atomic units and is the Wannier index (the value of $\overline{\mu}$ depends on the residual ion charge value, the numerical value of $\overline{\mu}$ for a unity charge of the residual ion is $\overline{\mu} = 1-127$). The interelectronic relative angle θ_{ab} is given by $\theta_{ab} := \cos^{-1} \hat{\mathbf{v}}_{a*}$ $\hat{\mathbf{v}}_b$. In case of higher excess energies (E \gg 1) it is readily verified that a -1 [Eq. (30)] and all modifications of the charges (23-25) which are due to incorporating the warnier treshold law become irrelevant. The charges (23-25) reduce then to those given in Ref. [9] with Eq. (21) being applied. From the functional forms of the charges (23-25) it is clear that when two particles approach each other (in velocity space) they experience their full two-body Coulomb interactions, whereas the third one 'sees' a net charge equal to the sum of the charges of the two close particles.

5. APPLICATIONS TO ATOMIC SCATTERING PROBLEMS

In this section we assess the analytical methods developed above by performing a

numerical evaluation of many-body scattering amplitudes. The reaction we are considering here is the electron-impact ionization of atomic hydrogen. In thenal channel of this collision process two interacting electrons move in the double continuum of a residual ion. Hence a correlated three-body wave function is needed to represent this state. For this wave function we employ the approximate expressions given in the preceding sections. The initial state consists of an incoming single-particle wave that represents the projectile electron and a bound state of atomic hydrogen.

The complete information on this reaction is obtained by measuring the coincidence rate for the emission of two continuum electrons with specified wave vectors, i.e. the energies E_a , E_b and the emission solid angles Ω_a , Ω_b of the two electrons are determined for a given incident energy of the projectile electron. Due to energy conservation it suffices to determine the energy of one of the electrons. Therefore, one measures in this way a triply differential cross section (TDCS), i.e. a cross section differential

 Ω_a , Ω_b and E_b .

If the spin of the electrons is not resolved, the TDCS is a statistically weighted average of singlet and triplet scattering cross sections

$$TDCS(\mathbf{k}_a, \mathbf{k}_b) = c\left(\frac{1}{4}|T^s|^2 + \frac{3}{4}|T^t|^2\right)$$
 (31)

where \mathbf{k}_i is the momentum of the incident projectile and $c = (2_{\pi})^4 (k_a k_b)/k_i$. The singlet T^s and triplet transition matrix elements T' derive from the corresponding transition operators T^s and T^t , where

$$\mathcal{T}^{s} = (\mathbf{I} + \mathcal{P}_{ab})\mathcal{T}_{fi}(\mathbf{k}_{a}, \mathbf{k}_{b})$$

$$\mathcal{T}^{t} = (\mathbf{I} - \mathcal{P}_{ab})\mathcal{T}_{fi}(\mathbf{k}_{a}, \mathbf{k}_{b}).$$
 (32)

The action of the exchange operator \mathcal{P}_{ab} on the operator \mathcal{T}_{fi} is given by $\mathcal{P}_{ab}\mathcal{T}_{fi}(\mathbf{k}_a, \mathbf{k}_b) = \mathcal{T}_{fi}$ (\mathbf{k}_b, k_a) . The prior representation of \mathcal{T}_{fi} (ka, kb) is given by

$$T_{fi}(\mathbf{k}_a, \mathbf{k}_b) = \langle \Psi | V_i | \Phi_{\mathbf{k}_i} \rangle . \tag{33}$$

The wave function Ψ is obtained from Eq. (18) as

$$\Psi_{\mathbf{k}_a,\mathbf{k}_a} = N \exp i(\mathbf{k}_a \cdot \mathbf{r}_a + \mathbf{k}_b \cdot \mathbf{r}_b) \bar{\Psi}_{\mathbf{k}_a,\mathbf{k}_a}.$$

The three-body system in the initial channel is described by $|\Phi_{ki}\rangle$. Assuming $|\Phi_{ki}\rangle$ to be the asymptotic initial-state, i. e. $\langle \mathbf{r}_a, \mathbf{r}_b | \Phi_{ki} \rangle$ is a product of an incoming plane wave representing the incident projectile electron and an undistorted 1s-state of atomic hydrogen, the perturbation operator V_i occurring in Eq. (33) is given by $1/|\mathbf{r}_a - \mathbf{r}_b| - 1/\mathbf{r}_a$ (which is the part of the total Hamiltonian not diagonalized by $|\Phi \mathbf{k}_i\rangle$). In what follows we choose the x axis as the incident direction $\hat{\mathbf{k}}_i$. The final state electrons are detected in a coplanar geometry, i. e. k_i . $(\mathbf{k}_a \times \mathbf{k}_b) = 0$. The z axis lies along the direction perpendicular to the scattering plane, i. e. parallel to $\hat{\mathbf{k}}_{\mathrm{a}} \times \hat{\mathbf{k}}_{\mathrm{b}}$. The polar and azimuthal angles of the vector k a (k b) are denoted by ka (kb) are denoted by θ_a , ϕ_a (θ_b , ϕ_b), respectively. In the coplanar geometry considered here the polar angles are fixed to $\theta_a = \pi/2 = \theta_b$. In the calculation of the DS3C model we employ the approximation (21) and use the product charges (23-25). If we use the unit matrix for the transformation (17), i.e. if we assume A = I, the three-body wave function reduces eigenfunction of the asymptotic part H_{par} of the Hamiltonian without any coupling to the internal region. This wave function is commonly known as the 3C wave function [3, 28]. In addition we compare the results of the analytical methods presented here with those of the convergent close coupling method (CCC). This is a purely numerical method that attempts at evaluating exactly the transition matrix elements fully numerically. In Fig.1 the angular distribution of one of the electrons is shown for two fixed angular positions of the other electron. The two electrons are ejected with equal energies $E_a = E_b$ = 6:8 eV . As clear from Fig. 1 the effect of the coupling to the internal region is very important, since the results of the 3C model that neglects

the short-range dynamics are at clear variance with the experiment. The differences between the CCC method and the experiments are still the subject of current research. The main advantage of analytical methods is that they allow an insight into the origin of the structures observed in the cross sections. An extensive analysis underlying this statement has been carried out in Ref.([31]) where the main peaks in Fig. 1 have been assigned to certain sequence of collisions between the participating particles.

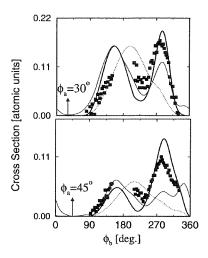


Figure 1: The fully differential cross section for the electron-impact ionization of atomic hydrogen in the co-planar, symmetric energysharing geometry. The incident energy is E_i = 27.2 eV. One ejected electron is detected at fixed angle Φ_a withe respect to the incident direction $[\Phi_a = 30^{\circ} \text{ (upper panel) and } \Phi_a =$ 45° (lower pannel)]. The angular distribution of the other emitted electron is measured. The emission angle of this electron with respect to the incident direction is denoted by Φ_b . Both electrons have the same energyn, namely $E_a =$ $E_b = 6.8 \text{ eV}$. Full squares are experimental data Ref. [30,31]. The solid thick lines show the predictions of the DS3C theory employing the matrix \mathcal{A} (cf. Eq.(17)) whereas the dotted curves indicated the results of the 3C theory, i.e. whin using $A \equiv 1$. Thin solid lines are the numerical calculations using convergent close coupling method (CCC).

At higher incident energies the discrepancies between the DS3C and the 3C results disappear and both of those models (as well as the CCC calculations) are in overall agreement with the experiments [36]. From this situation one can conclude that at higher energies the short-range parts of the Hamiltonian (H_{int} and H_{mix}) are of less importance, for they have been neglected in the 3C model whereas the DS3C theory accounts for them via the dynamical screening (we note the asymptotic region is reached for large $k_{ii} \xi_{k}$, i.e. for large momenta the distance ξ_k does not need to be very large).

This means in physical terms that the two electrons attains their asymptotic momenta swiftly without much of scattering from intermediate states whose behaviour is determined mainly by H_{int} and H_{mix} .

Integrating over all emission angles of the two electrons we end up with a single differential cross section depending on the energy of one of the electrons. Since the energy of the other electron is then determined via the energy-conservation law, the single differential cross section has to be symmetric with respect to the point where both electrons have the same energy. Fig. 2 shows the results for the single differential cross sections as calculated within the DS3C method along with the calculations within the 3C method. The excess energy is very low (200 meV).

For small excess energies the Wannier theory, which relies on phase space arguments, predicts a flat energy distribution between the electrons, i.e. a flat single-differential cross section. This prediction has been substantiated by full numerical calculations [37]. As seen in Fig. 2 the DS3C predicts a flat energy sharing between the electrons close to the complete fragmentation threshold, in contrast to the 3C results which are strongly peaked around the energy-sharing configuration. deviation of the 3C results from those of the Wannier theory is not surprising since in the Wannier approach one expands the potential around a saddle point (accounting for terms up to a fourth order) and neglects higher order

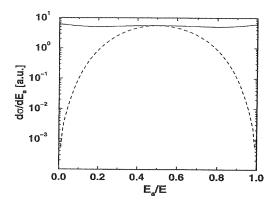


Figure 2: The single differential cross section for the electron-impact ionization of atomic hydrogen as function of the ratio E_a/E where E_a is the energy of one of the final-state electrons and $E = E_a + E_b$ is the total excess energy which is chosen as E = 200 meV. The use of the Ψ_{DS3C} approximation yields the solid curve whereas the dashed curve represents the results when employing the 3C model ($\mathcal{A} = I$). The 3C results have been multiplied by a factor of 10^{13} for a better shape comparison.

terms while the kinetic energy is treated fully. In contrast the 3C model neglects the short-range part of the kinetic energy. Obviously it is this part which is most important for the Wannier mode and the resulting predictions.

Sampling over the energy sharing between the two electrons, i.e. integrating the single differential cross section shown in Fig. 2, one obtains the total cross section as function of the excess energy $E = E_a + E_b$ (or equivalently as function of the incident energy E_i). Close to the three-body break-up threshold the total cross section $\sigma(E)$ for two continuum electrons receding from a charged ion has been investigated by Wannier [17] using a classical analysis. Wannier [17] pointed out that the excess-energy functional dependence of the total ionization cross section at the threeparticle fragmentation threshold can deduced from the volume of the phase space available for double escape of the two electrons. For the present case of atomic hydrogen Wannier deduced the threshold law

 $\sigma(E) \propto E^{1.127}$. Since then an immense amount of theoretical and experimental studies (e.g. [18-25,38-41]) using quite different approaches have been carried out which basically confirm the Wannier-threshold law.

The Wannier treatment predicts the scaling behaviour of the cross section $\sigma(E)$, but it does not provide any information about the magnitude of $\sigma(E)$. That the magnitude is a very sensitive quantity is illustrated by the behaviour of the cross section in the independent Coulomb particle model which is obtained our case by switching the interaction between the two electrons in the final channel. In this case the cross section reveals a linear dependence on the excess energy, $\sigma(E) \propto E$ [42]. Although the latter dependence of $\sigma(E)$ does not deviate much from the Wannier threshold law ($\sigma(E) \propto E^{1.127}$) the absolute value of $\sigma(E)$ within the inde-pendent Coulomb particle model is largely overestimated [compare Fig.3]. If we employ the wave function Ψ_{DS3C} , with the dynamical product charges described in the preceding sections we end up with results in good accord with the experimental measurements (cf. Fig.3). The absolute magnitude of the total cross section is satisfactorily reproduced when the DS3C model is employed. To examine the analytical behaviour of $\sigma(E)$ calculated using Ψ_{DS3C} we plot in Fig. 3(b) the quantity $\sigma(E)/E^{1.127}$. According to the Wannier-threshold law the latter quantity should be a constant function of E and gives the absolute value of the cross section. As seen in Fig. 3(b) the Wannier threshold law is in fact reproduced by the cross section results of the Ψ_{DS3C} within a range of E [0, 0.5eV]. For E > 0.5 eV the analytical dependence of $\sigma(E)$ evaluated with Ψ_{DS3C} slowly deviates from the Wannier threshold law. When using the 3C method for the description of the two escaping electrons (A = I in Eq.(17)) we obtain an analytical behaviour for $\sigma(E)$ which is not compatible with the Wannier theory. The absolute value for the total cross section is as well not reproduced by the 3C model, for the reasons discussed above. Also included in Fig.3

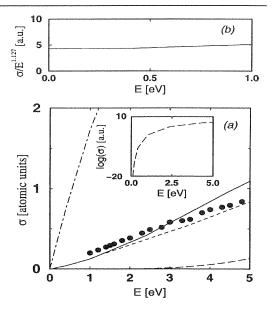


Figure 3: The total cross section $\sigma(E)$ for the electron-impact ionization of atomic hydrogen as a function of the excess energy E. The solid (long dashed) curve shows the results for $\sigma(E)$ when treating the DS3C theory (3C model) whereas the dashed dotted curve denotes the results of the independent Coulomb particles model (see text). Results of the CCC method are also included (short dashed curve). Experimental data are due to Shah $et\ al.\ [32]$. Te inset in the panel (a) shows the results of the 3C theory on a logarithmic scale. In the upper panel (b) the quantity $\sigma(E)/E^{1.127}$ is depicted as a function of E as evaluated using Ψ_{DS3C} .

are the results of the convergent-close coupling method, CCC, [33]. The results of the CCC are in good agreement with the experimental $\sigma(E)$ for higher energies [33], however, close to threshold the evaluation of $\sigma(E)$ is limited by the computational resources as an ever increasing number of pseudo states is needed to achieve convergence.

In addition to the magnitude of the cross sec-tion, the spin asymmetry, A, offers a further way of probing the dynamical properties of the electron-impact ionization of atomic systems. The spin asymmetry A is defined as

$$A(E) := \frac{\sigma^s(E) - \sigma^t(E)}{\sigma^s(E) + 3\sigma^t(E)},\tag{34}$$

where σ^{s} and σ^{t} are the total ionization cross sections for singlet and triplet scattering, respectively. The Wannier theory for threshold ionization predicts a constant value of A with increasing excess energy but provides no information on the numerical value of A [43]. Measurements of A at threshold reveals a slightly positive slope of the spin asymmetry with increasing excess energies [44]. In Fig.4 the results for A are shown in the case where the two-electron continuum final state is described by the 3C theory and by Ψ_{DS3C} . Also depicted in Fig.4 are the results of the CCC approach [33] and the method using hiddencrossing theory [25]. Although all theories, except for the 3C model, are in reasonable agreement with experimental finding the positive slope of A at threshold is not reproduced.

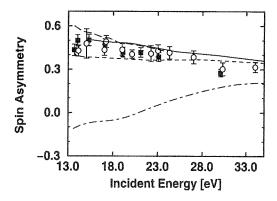


Figure 4: The spin asymmetry [Eq.(34)] in the total ionization cross section $\sigma(E)$ for the electron-impact ionization of atomic hydrogen. Results of the hidden crossing theory [25] (long dashed curve) and those of the CCC method [33] (solid curve) are depicted. Treating the final stated according to the DS3C theory (3C theory) yields for the spin asymmetry the results shown by the short-dashed (dot-dashed) curve. The experimental data are due to Fletcher *et al.* [34] (full squares) and Crowe *et al.* [35] (open circles).

Neglect of the short-ranged part of the Hamiltonian H_{in} and H_{mix} , i.e. using the 3C model, results in a completely wrong behaviour of the calculated spin asymmetry. With increasing excess energy the inner region of the Hilbert space becomes of less importance for the present reaction and the results of the 3C method become more and more in better agreement with the experimental data.

We note here that since the spin asymmetry is a ratio of cross sections it is expected that the spin asymmetry is rather sensitive to the detailed of the radial part of the wave functions. From the agreement between the experiment and the DS3C theory observed in Fig. 4 we conclude that the radial part of $\Psi_{\rm DS3C}$ is well behaved at lower excess energies and that the short-range parts of the total Hamiltonian H_{int} and H_{mix} plays a dominate role at lower energies, as far as the value and the behaviour of the spin asymmetry are concerned.

6. GREEN FUNCTION THEORY OF FINITE SYSTEMS

In the preceding sections we investigated the two and three-body problem. With increasing number of particles the treatment becomes more complex and a methodology different from the wave function technique is more appropriate. A method which is widely used in theoretical physics is the Green function approach which we will followup in this section.

For a canonical ensemble, we seek a non-perturbative method which allows to distribute systematically the total energy between the potential and the kinetic energy parts. This is achieved by the development of an incremental method in which the N correlated particle system is mapped exactly onto a set of systems in which only N-M particles are interacting $(M \in [1, N-2])$, i.e. in which the potential energy part is damped. (In contrast to renormalization group theory we do not reduce the strength of interactions, but the number of them). This is particularly interesting from a

thermodynamic point of view since for a number of thermodynamic properties the kinetic energy contributions can be separated out from the potential energy parts, as shown in the next section for the internal energy. By virtue of the present method the potential energy part is systematically reduced.

For a formal development let us consider a nonrelativistic system consisting of Ninteracting particles. We assume the total potential to be of the class $U^{(N)} = \sum_{i>i=1}^{N} v_{ij}$ without any further specification of the individual potentials v_{ij} . For three-body potentials v_{ij} . For three-body potentials the development of the theory poceeds along the same lines.

The potential $U^{(N)}$ satisfies the recurrence relations

$$U^{(N)} = \frac{1}{N-2} \sum_{j=1}^{N} u_j^{(N-1)}, \tag{35}$$

$$u_j^{(N-1)} = \frac{1}{N-3} \sum_{k=1}^{N-1} u_{jk}^{(N-2)}, \ j \neq k, \ (36)$$

where $u_i^{(N-1)}$ is the total potential of a system of N-1 interacting particles in which the j particle is missing, i.e in terms of the physical pair potentials v_{mn} , one can write $u_i^{(N-1)}$

$$= \ddagger^{N}_{m > n = 1} v_{mn}, \quad m \neq j \neq n.$$

The fundamental quantity that describes the microscopic properties of the N body quantum system is the Green operator $G^{(N)}$ which is the resolvent of the total Hamiltonian. It can be deduced from the Lippmann Schwinger equation $G^{(N)} = G_0 + G_0 U^{(N)} G^{(N)}$ where G_0 is the Green operator of the non interacting N body system. An equivalent approach to determine the dynamical behavior of a system is to derive the respective transition operator $T^{(N)}$ which satisfies the integral equation $T^{(N)} = U^{(N)} + U^{(N)} G_0 T^{-(N)}$. These integral equations for $G^{(N)}$ and $T^{(N)}$ provide a natural framework for perturbative treatments.

However, for $N \ge 3$ the application of the above Lippmann Schwinger equations (and those for the state vectors) is hampered by mainly two difficulties: 1.) as shown in Refs. [45,46] the Lippmann Schwinger equations for the state vectors do not have a unique solution, and 2.) as shown by Faddeev [47-49] the kernel of these integral equations $K = G_0 U^{(N)}$ is not a square integrable operator for $N \ge 3$, i.e. the norm $||K|| = [Tr(KK^{\dagger})]^{1/2}$ is not square integrable. The kernel K is also not compact. The reason for this drawback is the occurrence of the so-called disconnected diagrams where one of the N particles is a spectator, i.e. not correlated with the other N-1 particles. For the three-body problem Faddeev [47,48] suggested alternative integral equations with square integrable kernel.

Our aim here is twofold: (a) We would like to derive non-perturbative integral equations that treat all N particles on equal footing and are free from disconnected diagrams. (b) These equations should allow to obtain, in a computationally accessible manner, solution of the correlated N body problem from the solution when only N - M particles are interacting (where $M \in [1, N-2]$).

According to the decomposition (35), the intergral equation for the transition operator can be written as

$$T^{(N)} = \sum_{i=1}^{N} T_i^{(N-1)} \tag{37}$$

$$T_j^{(N-1)} = \tilde{u}_j^{(N-1)} + T^{(N)} G_0 \tilde{u}_j^{(N-1)}, \ j \in [1, N].$$
 (38)

Here we introduced the scaled potential

 $\tilde{u}_i^{(N-1)} = \left(u_i^{(N-1)}\right)/(N-2).$

The transition operator of the system, when N-1 particles are interacting via the scaled potential $\bar{u}_{j}^{(N-1)}$, is

$$t_j^{(N-1)} = \tilde{u}_j^{(N-1)} + \tilde{u}_j^{(N-1)} G_0 t_j^{(N-1)}.$$

With this relation Eq. (38) can be reformulated as

$$T_{j}^{(N-1)} = t_{j}^{(N-1)} + t_{j}^{(N-1)} G_{0} T^{(N)} - t_{j}^{(N-1)} G_{0} \left(\tilde{u}_{j}^{(N-1)} + \tilde{u}_{j}^{(N-1)} G_{0} T^{(N)} \right)$$

$$= t_{j}^{(N-1)} + t_{j}^{(N-1)} G_{0} \left(T^{(N)} - T_{j}^{(N-1)} \right)$$

$$= t_{j}^{(N-1)} + t_{j}^{(N-1)} G_{0} \sum_{k \neq j}^{N} T_{k}^{(N-1)}. \tag{39}$$

Eq.(39) can be expressed in a matrix form as follows

$$\begin{pmatrix} T_{1}^{(N-1)} \\ T_{2}^{(N-1)} \\ \vdots \\ T_{N-1}^{(N-1)} \\ T_{N}^{(N-1)} \end{pmatrix} = \begin{pmatrix} t_{1}^{(N-1)} \\ t_{2}^{(N-1)} \\ \vdots \\ t_{N-1}^{(N-1)} \\ t_{N}^{(N-1)} \end{pmatrix} + [\mathbf{K}^{(N-1)}] \begin{pmatrix} T_{1}^{(N-1)} \\ T_{2}^{(N-1)} \\ \vdots \\ T_{N-1}^{(N-1)} \\ T_{N}^{(N-1)} \end{pmatrix}$$

$$(40)$$

The kernel $[\mathbf{K}^{(N-1)}]$ is a matrix operator and is given by

$$\begin{bmatrix} \mathbf{K}^{(N-1)} \end{bmatrix} = \begin{bmatrix} 0 & t_1^{(N-1)} & t_1^{(N-1)} & \dots & t_1^{(N-1)} \\ t_2^{(N-1)} & 0 & t_2^{(N-1)} & \dots & t_2^{(N-1)} \\ \dots & \dots & \dots & \dots & \dots \\ t_{N-1}^{(N-1)} & \dots & t_{N-1}^{(N-1)} & 0 & t_{N-1}^{(N-1)} \\ t_N^{(N-1)} & \dots & t_N^{(N-1)} & t_N^{(N-1)} & 0 \end{bmatrix} G_0.$$
 (41)

From Eq. (36) it is clear that $t_j^{(N-1)}$ can also be expressed in terms of the transition operators of the system where only N-2 particles are interacting:

$$t_j^{(N-1)} = \sum_{k \neq j}^{N-1} T_k^{(N-2)}.$$

The operators $T_k^{(N-2)}$ are deduced from Eq. (40) with N being replaced by N-1.

From the relation $G^{(N)} = G_0 + G_0 T^{(N)} G_0$ we conclude that the Green operator of the interacting N particle system has the form

$$G^{(N)} = G_0 + \sum_{j=1}^{N} G_j^{(N-1)}.$$
 (42)

The operators $g_j^{(N-1)}$ are related to the Green operators $g_j^{(N-1)}$ of the systems in which only N-1 particles are correlated by virtue of $\tilde{u}_j^{(N-1)}$. This interrelation is given via

$$\begin{pmatrix}
G_{1}^{(N-1)} \\
G_{2}^{(N-1)} \\
\vdots \\
G_{N-1}^{(N-1)} \\
G_{N}^{(N-1)}
\end{pmatrix} = \begin{pmatrix}
g_{1}^{(N-1)} - G_{0} \\
g_{2}^{(N-1)} - G_{0} \\
\vdots \\
g_{N-1}^{(N-1)} - G_{0} \\
g_{N}^{(N-1)} - G_{0}
\end{pmatrix} + [\tilde{\mathbf{K}}^{(N-1)}] \begin{pmatrix}
G_{1}^{(N-1)} \\
G_{2}^{(N-1)} \\
\vdots \\
G_{N-1}^{(N-1)} \\
G_{N}^{(N-1)}
\end{pmatrix}, (43)$$

where $[\tilde{\mathbf{K}}^{(N-1)}] = G_0[\tilde{\mathbf{K}}^{(N-1)}]G_0^{-1}$. From Eqs. (40,43) we conclude that if the Green operator of the interacting N-1 body system is known (from other analytical or numerical procedures, e.g. from an effective field method. such as density functional theory) the Green operator of the N particles can then be deduced by solving a set of N particles can then be deduced by solving a set of N linear, coupled integral equations (namely Eqs. (40,43)). According to the above equations (namely Eqs. (40,43)). According to the above equations, if only the solution of the N-M problem is known where $M \in [1, N-2]$ we have to perform a hierarchy of calculations starting by obtaining the solution for te N- M+1 problem and repeating the procedure to reach the solution of the *N* body problem.

At first sight the kernels of Eqs. (40,43) appear to have disconnected diagrams since they contain transition operators of systems where only N-1 particles are interacting and one particle is free (disconnected). It is, however, straightforward to show that any iteration of these kernels is free of

disconnected terms (the disconnected terms occurs only in the off-diagonal elements of $[\mathbf{K}^{N-M}]$ and $[\tilde{\mathbf{K}}^{N-M}]$). For N=3 the present reduces to the well-established scheme Faddeev equations. As for the functional structure of the Eqs. (40,43) we remark that for the solution of the N particle problem we need the (off-shell) transition operators of the N-1sub-system. The interaction potentials do not appear in this formulation (in contrast to the Lippmann Schwinger approach). On the other hand the (on-shell) transition matrix elements can be determined experimentally, This fact becomes valuable when the potentials are not known.

6.1. APPLICATION TO FOUR-BODY SYSTEMS

Over the years a substantial body of knowledge on the three-particle problem has been accumulated. In contrast, theoretical studies on the four-body problem are still scare due to computational limitations whereas an impressive amount of experimental data is already available [50{54}]. Thus, it is desirable to apply the above procedure to the four-body system and to express its solution in terms of known solutions of the three body problem. For N = 4 the first iteration of Eq.(43) yields

$$G^{(4)} = \sum_{j=1}^{4} g_j^{(3)} - 3G_0. \tag{44}$$

Here $g_j^{(3)}$ is the Green operator of the system where only three particles are interacting and can be taken from other numerical or analytical studies. This means, to a first order, methods treating the correlated three-body problem can be extended to deal with the four-body case using Eq.(44). We note that for the case of non-interacting system $g_j^{(3)}$ reduces to $g_j^{(3)} \equiv G_0$ and hence Eq. (44) reduces to $G^{(4)} = G_0$, as expected.

The Green function encompasses the complete spectrum of the many-body system, i.e. the wave function approach can be

retrieved from the Green function. For example, Eq. (44) leads to an expression for the four-body state vector in the form

$$|\Psi^{(4)}\rangle = |\psi_{234}^{(3)}\rangle + |\psi_{134}^{(3)}\rangle + |\psi_{124}^{(3)}\rangle + |\psi_{123}^{(3)}\rangle -3|\phi_{\text{free}}^{(4)}\rangle.$$
(45)

Here $|\psi_{ijk}^{(3)}\rangle$ is the state vector of the system in which the three particles i, j and k are interacting whereas $|\phi_{\text{free}}^{(4)}\rangle$ is the state vector of the non-interacting four-body system. The state vectors $|\psi_{ijk}^{(3)}\rangle$ can be approximated by Eq. (18) or by the other procedures discussed in the preceding section on the three-body problem.

Since the state vector (45) is expressed as a sum of correlated three-body states, the evaluation of the four-body transition matrix elements for a specific reaction simplifies considerably. In addition, the properties of a many-body interacting system can be obtained in a straightforward way from those for systems with a reduced number of interactions, for in this case the matrix elements of the total Green functions are expressed as sums of matrix elements of reduced Green functions, as evident from Eq.(44). This spectral feature can be exploited to study the thermodynamical properties of finite correlated systems.

7. THERMODYNAMICS AND PHASE TRANSITIONS OF INTERACTING FINITE SYSTEMS

To investigate the thermodynamical properties of N interacting particle system we remark that at the critical point divergent thermodynamical quantities, such as the specific heat C_V are obtained as a derivative with respect to the inverse temperature β of the logarithm of the canonical partition function $Z(\beta)$,

$$C_V = \beta^2 \partial_{\beta}^2 \ln Z(\beta) = f(\beta, Z(\beta))/Z(\beta).$$

Here f is some analytical function and for the Boltzmann constant we assume k = 1.

Therefore divergences in the thermodynamic quantities, which signify phase transitions are connected to the zero points of $Z(\beta)$. These zero points are generally complex valued. Therefore an analytical continuation of $Z(\beta)$ to complex temperatures is needed.

The connection between the phase transitions and the complex zero points of the grand canonical partition function have been uncovered by Yang and Lee [55]. In this case one seeks an analytical continuation of the fugacity $z=\exp(\beta\mu)$ (here μ is the chemical potential) to the complex plane $z\to\Re(z)+i\Im(z)$. In the thermodynamic limit the zero points condense to lines. The transition points are the crossing points of these lines with the real fugacity axis.

Grossman *et al.* [56] generalized the concept of Yang and Lee to the canonical ensemble. In this case the inverse temperature $\beta = 1/T$ is continued analytically to $\beta = \Re(\beta) + i\Im(\beta)$. The phase transitions are then the crossings of the zero points line of $Z(\beta)$ with the real β axis. The advantage here is that a classification of the phase transitions can be given in terms of how the zero-points line do cross the real β axis [56].

The crucial point is that in the thermodynamic limit $N \to \infty$, $V \to \infty$ and v = V = N < ∞ (V is the volume) the zero points approach, to an in infinitesimal small distance the real axis. For this reason, the characteristic phasetransition divergences appear thermodynamical quantities. For finite systems $Z(\beta)$ has only finite zero points that do not approach infinitely close the real axis. Therefore, the thermodynamic quantities show smooth peaks rather then divergences. The positions and widths of these peaks can be obtained from the real and imaginary parts of the zero points laying closest to the real axis [57].

To apply this method to correlated finite systems we need a representation of the canonical partition function that can then be continued analytically to the complex temperature plane.

7.1. REPRESENTATION OF THE CANONICAL PARTITION FUNCTION

The canonical partition function of a correlated system can be expressed in terms of the many-body Green function as

$$Z(\beta) = \int dE \ \Omega(E) \ e^{-\beta E}. \tag{46}$$

Here $\Omega(E)$ is the density of states which is related to imaginary part of the trace of $G^{(N)}$ via

$$\Omega(E) = -\frac{1}{\pi} \Im \operatorname{Tr} G^{(N)}(E). \tag{47}$$

From the Green function expansion Eq.(42) we deduce

$$Z(\beta) = -\frac{1}{\pi} \Im \int dE \operatorname{Tr} G^{(N)}(E) e^{-\beta E}$$
$$= Z_0(\beta) + \sum_{i=1}^{N} Z_i(\beta)$$
(48)

where

$$Z_0(\beta) = -\frac{1}{\pi} \Im \int dE \, \text{Tr} \, G_0(E) \, e^{-\beta E}$$
 (49)

$$Z_j(\beta) = -\frac{1}{\pi} \Im \int dE \operatorname{Tr} G_j^{(N-1)}(E) e^{-\beta E}.$$
 (50)

To a first order $z_j = z_j^{(N-1)} - z_0$ where $z_j^{(N-1)}$ is the partition function of a system in which only N-1 particles are interacting. For the applications of the Grossmann method let us remark that $Z(\beta)$ is an integral function and can be expressed in a polynomial form. Recalling the analytical properties of meromorphic functions one can write $Z(\beta)$ in terms of its complex zero points as

$$Z(\beta) = Z(0)e^{\beta \frac{Z'(0)}{Z(0)}} \prod_{k=1}^{\infty} \left(1 - \frac{\beta}{\beta_k}\right) e^{\frac{\beta}{\beta_k}}.$$
 (51)

7.1.1. APPLICATIONS: THE INTERNAL ENERGY AND THE SPECIFIC HEAT

As an application let us consider the internal energy and the specific heat of an interacting system. The internal energy is

defined as $U = -\partial_{\beta} \operatorname{In} Z(\beta)$. From Eq. (51) we deduce

$$U = -\partial_{\beta} \ln \left[Z(0) e^{\beta \frac{Z'(0)}{Z(0)}} \prod_{k=1}^{\infty} \left(1 - \frac{\beta}{\beta_k} \right) e^{\frac{\beta}{\beta_k}} \right]$$

$$= -\frac{Z'(0)}{Z(0)} - \sum_{k} \partial_{\beta} \left[\ln \left(1 - \frac{\beta}{\beta_k} \right) \right] + \frac{1}{\beta_k}$$

$$= -\frac{Z'(0)}{Z(0)} + \sum_{k} \frac{1}{\beta_k - \beta} - \frac{1}{\beta_k}.$$
 (52)

The part $-\frac{z'(0)}{z(0)}$ corresponds to the

internal energy of the ideal gas
$$\left(-\frac{z'(0)}{z(0)} = \frac{3N}{2\beta}\right)$$
,

since for $\beta \to 0$ $(T \to \infty)$ the kinetic energy dominates over any other interaction energy in the system and the internal energy admits thus the form of the ideal gas internal energy. Eq. (52) means that in a finite (or extended system) it is always possible to separate out the kinetic energy part. This part is smooth and has no zero points. The potential energy will thus be the dominant factor as far as the zero points of the partition function are concerned. The zero of complex points of Z occur as pairs $[\beta_k, \beta_k^*] =$ conjugates. $[\Re(\beta_k) +$ $i\Im(\beta_k)$, $\Re(\beta_k) - i\Im(\beta_k)$]. The internal energy therefore be written as

$$U = \frac{3N}{2} + \sum_{k} \frac{1}{\beta_{k} - \beta} + \frac{1}{\beta_{k}^{*} - \beta}$$

$$-\frac{1}{\beta_{k}} - \frac{1}{\beta_{k}^{*}}$$

$$= \frac{3N}{2} + 2\sum_{k} \frac{\Re(\beta_{k}) - \beta}{(\Re(\beta_{k}) - \beta)^{2} + (\Im(\beta_{k}))^{2}}$$

$$-\frac{\Re(\beta_{k})}{|\beta_{k}|^{2}}.$$
(53)

For the specific heat we arrive at

$$C_V = -\beta^2 \partial_{\beta} U = \beta^2 \partial_{\beta}^2 \ln(Z(\beta))$$
$$= \frac{3N}{2} + \beta^2 \sum_k \partial_{\beta} \left(\frac{1}{\beta_k^{-\beta}} \right)$$

$$= \frac{3N}{2} + \sum_{k} \left(\frac{\beta}{\beta_{k}^{-}\beta}\right)^{2}$$

$$= \frac{3N}{2} -$$

$$2\sum_{k} \frac{\left[\Re(\beta_{k}) - \beta\right]^{2} - \left[\Im(\beta_{k})\right]^{2}}{\left(\left[\Re(\beta_{k}) - \beta\right]^{2} + \left[\Im(\beta_{k})\right]^{2}\right)^{2}}.$$
(54)

Now we can employ the formulas Eqs. (46-50) to derive the zero points which enter Eqs. (53,54). In this context it is important to recall that the reduction procedure presented here expresses the total interacting Green function in terms of systems with the same number of particles but with less number of interactions. That means in each reduction cycle of G(N) the interaction energy is reduced while the kinetic energy is kept constant.

8. CONCLUSIONS

In this work an overview was given on the recent development in the analytical treatment of correlated few charged-particle systems. For a three-body system we discussed the existence of an approximate separability of the long and short-range dynamics. This separability is exposed by introducing curvilinear coordinates in which case the total Hamiltonian of the system is expressed as a sum of two commuting operators: A short range operator that decays for large inter-particle distances and a long-range operator that possesses an oscillatory behaviour in the asymptotic region. At finite distances the total Hamiltonian contains in addition to these two operators a term which mixes the long and the short range dynamics.

Approximate many-body wave functions in the asymptotic region and at finite distances have been derived and their range of validity have been discussed. The strengths and shortcomings of the derived wave functions are exposed by evaluating scattering transition matrix elements and comparing to available experimental data. In a further section we introduced a Green function methodology that is capable of dealing with large finite interacting systems. numerical Α implementation of this method for the

calculations of scattering amplitudes can be found in Ref. [58] whereas an extension to include ordered and disordered scattering potentials can be found in Ref. [59]. As a further application of this method we discussed briefly the thermodynamic properties of interacting finite systems and derived an expression for the canonical partition function.

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