Spectral Density and Metal-Insulator Phase Transition in Mott Insulators within Reduced Density Matrix Functional Theory

S. Sharma,^{1,*} J. K. Dewhurst,¹ S. Shallcross,² and E. K. U. Gross¹

¹Max-Planck-Institut für Mikrostrukturphysik, Weinberg 2, D-06120 Halle, Germany

²Lehrstuhl für Theoretische Festkörperphysik, Staudstrasse 7-B2, 91058 Erlangen, Germany

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We present a method for calculating the spectrum of periodic solids within reduced density matrix functional theory. This method is validated by a detailed comparison of the angular momentum projected spectral density with that of well-established many-body techniques, finding very good agreement in all cases. The physics behind the pressure induced insulator-metal phase transition in MnO is investigated. The driving mechanism of this transition is identified as increased crystal field splitting with pressure, resulting in a charge redistribution between the Mn e_g and t_2g symmetry projected states.

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Transition metal oxides (TMOs), the prototypical Mott insulators, are test-bed systems for new functionals within density functional theory (DFT) and many-body theories alike. Spectra at zero temperature and pressure obtained from many-body theories are in good agreement with experiments. Moreover, the spectral density obtained using dynamical mean-field theory (DMFT) [1] and the G_0W_0 corrected DFT [2] agree with each other even for subtle features such as symmetry and site projected spectral density. Single particle DFT spectra can also be made to agree with these many-body results by using two separate fitting parameters: the on-site Coulomb term U and the scissors shift Δ , where Δ is the difference between the experimental gap and the Kohn-Sham gap obtained using the local spin density approximation + U functional [2].

Away from the ground state, TMOs show the rich physics of insulator-metal phase transitions. The classic Mott insulator, MnO, exhibits metallization under pressure. This phase transition is accompanied by a simultaneous moment and volume collapse [3–8]. On the theory side, however, the physics of this phase transition is totally different for different methods; while DFT results indicate that the increase in bandwidth controls the phase transition [9,10], DMFT results, on the other hand, show that the main reason for metallization lies in the increased crystal field splitting [1].

Recently, reduced density matrix functional theory (RDMFT) has shown potential for correctly treating Mott insulators under ambient conditions [11,12]. RDMFT is an appealing alternative because it does not require any system dependent parameters and thus is a truly *ab initio* theory for treating strong correlations. However, it still remains to be seen how RDMFT performs away from ambient pressure conditions; can RDMFT capture the insulator-metal phase transition? What is the physics of this phase transition within RDMFT? In order to answer these questions one requires two things: (1) a magnetic extension of RDMFT and (2) information about the photoemission spectrum to shed light on the nature of the phase transition. The latter

is a difficult quantity to extract from RDMFT, which, by its very nature, is a ground-state theory.

In the present work we extend RDMFT to describe magnetic solids and further present a technique for calculating the photoemission spectrum. We validate this technique by demonstrating the agreement of the t_{2g} and e_g resolved spectral density thus obtained, with the wellestablished many-body methods like *GW* and DMFT. We further show that not only at ambient pressure but also away from it RDMFT correctly determines the spectra of Mott insulators and captures the physics of the insulator to metal phase transition.

Within RDMFT, the one-body reduced density matrix (1-RDM) is the basic variable [13,14]

$$\gamma(\mathbf{x}, \mathbf{x}') \equiv N \int d^3 x_2 \dots d^3 x_N \Psi(\mathbf{x}, \mathbf{x}_2 \dots \mathbf{x}_N) \Psi^*(\mathbf{x}', \mathbf{x}_2 \dots \mathbf{x}_N),$$
(1)

where Ψ denotes the many-body wave function, N is the total number of electrons, and $\mathbf{x} = \{\mathbf{r}, \sigma\}$. Diagonalization of γ produces a set of orthonormal Bloch functions, the so-called natural orbitals [13] $\phi_{i\mathbf{k}}$, and occupation numbers $n_{i\mathbf{k}}$. In the present work we have extended RDMFT to the truly noncollinear magnetic case by treating the natural orbitals as two component Pauli spinors, leading to the spectral representation: $\gamma(\mathbf{x}, \mathbf{x}') = \sum_{i\mathbf{k}} n_{i\mathbf{k}} \varphi_{i\mathbf{k}}(\mathbf{r}) \otimes \varphi_{i\mathbf{k}}^{\dagger}(\mathbf{r}')$ with $\varphi_{i\mathbf{k}}(\mathbf{r}) \equiv \{\phi_{i\mathbf{k}}^{\dagger}(\mathbf{r}), \phi_{i\mathbf{k}}^{\dagger}(\mathbf{r})\}$. The necessary and sufficient conditions for ensemble N representability of γ [15] require $0 \le n_{i\mathbf{k}} \le 1$ for all i and \mathbf{k} , and $\sum_{i\mathbf{k}} n_{i\mathbf{k}} = N$.

In terms of γ , the total ground-state energy [14] of the interacting system is (atomic units are used throughout)

$$E[\boldsymbol{\gamma}] = -\frac{1}{2} \operatorname{tr}_{\sigma} \int \lim_{\mathbf{r} \to \mathbf{r}'} \nabla_{\mathbf{r}}^{2} \boldsymbol{\gamma}(\mathbf{x}, \mathbf{x}') d^{3} r' + \int \rho(\mathbf{r}) V_{\text{ext}}(\mathbf{r}) d^{3} r$$
$$+ \frac{1}{2} \int \frac{\rho(\mathbf{r}) \rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} d^{3} r d^{3} r' + E_{\text{xc}}[\boldsymbol{\gamma}], \qquad (2)$$

0031-9007/13/110(11)/116403(5) TH-2013-06 where $\rho(\mathbf{r}) = \text{tr}_{\sigma} \gamma(\mathbf{x}, \mathbf{x})$, V_{ext} is a given external potential, and $E_{\rm xc}$ we call the exchange-correlation (xc) energy functional. In principle, Gilbert's [14] generalization of the Hohenberg-Kohn theorem to the 1-RDM guarantees the existence of a functional $E[\gamma]$ whose minimum, for fixed a V_{ext} , yields the exact γ and the exact ground-state energy. In practice, however, the correlation energy is an unknown functional of γ and needs to be approximated. While there are several known approximations for the xc energy functional, the most promising for extended systems is the power functional [11] where the xc energy is given by $E_{\rm xc}[\gamma] = -1/2 \iint d^3x' d^3x |\gamma^{\alpha}(\mathbf{x}, \mathbf{x}')|^2 / |\mathbf{r} - \mathbf{r}'|$ with α indicating the power in the operator sense. In view of the universality of the functional $E_{\rm xc}[\gamma]$, the value of α should, in principle, be system independent. A few optimum values of α have been suggested in the literature [11,16,17]. In the present work α is fixed to 0.656 for all materials studied.

In order to devise a theoretical method to approximately obtain the spectral density we start from the definition of the retarded Green's function written in the basis of the natural orbitals

$$iG^{R}_{\lambda\lambda'}(t-t') = \Theta(t-t')\langle \Psi^{N}_{0}|\{a_{\lambda}(t), a^{\dagger}_{\lambda'}(t')\}|\Psi^{N}_{0}\rangle, \quad (3)$$

where $\lambda \equiv \{i, \mathbf{k}\}$ with the index *i* labeling the natural orbitals for a given \mathbf{k} . *a*, a^{\dagger} are the creation and annihilation operators associated with the complete set of natural orbitals and $|\Psi_0^N\rangle$ is the neutral *N*-electron ground state. The spectral function $A_{\lambda\lambda'}(\omega)$ can be written in terms of the Lehmann representation as

$$A_{\lambda\lambda'}(\omega) = -2\Im G^{\mathsf{R}}_{\lambda\lambda'}(\omega)$$

= $2\pi \sum_{j} \langle \Psi^{N}_{0} | a_{\lambda} | \Psi^{N+1}_{j} \rangle \langle \Psi^{N+1}_{j} | a^{\dagger}_{\lambda'} | \Psi^{N}_{0} \rangle$
 $\times \delta(\omega - [E^{N+1}_{j} - E^{N}_{0}]) + 2\pi \sum_{i} \langle \Psi^{N}_{0} | a^{\dagger}_{\lambda'} | \Psi^{N-1}_{i} \rangle$
 $\times \langle \Psi^{N-1}_{i} | a_{\lambda} | \Psi^{N}_{0} \rangle \delta(\omega - [E^{N}_{0} - E^{N-1}_{i}]), \qquad (4)$

where $H|\Psi_i^{N\pm 1}\rangle = E_i^{N\pm 1}|\Psi_i^{N\pm 1}\rangle$ is satisfied by the exact $(N \pm 1)$ -particle eigenstates of the Hamiltonian *H*. To deduce an approximate expression for the spectral function, we replace the complete set of eigenfunctions $\{|\Psi_i^{N\pm 1}\rangle\}$ by the set of approximate eigenfunctions obtained by adding (removing) a single electron in a natural orbital to (from) the exact correlated *N*-particle ground state:

$$\begin{split} |\Phi_{\zeta}^{N+1}\rangle &= \frac{1}{\sqrt{1 - n_{\zeta}}} a_{\zeta}^{\dagger} |\Psi_{0}^{N}\rangle, \\ |\Phi_{\zeta}^{N-1}\rangle &= \frac{1}{\sqrt{(n_{\nu})}} a_{\zeta} |\Psi_{0}^{N}\rangle. \end{split}$$
(5)

While these many-body states are clearly not a complete set, we do expect them to capture the dominant contributions to direct and inverse photoemission. The set $\{|\Phi_i^{N\pm 1}\rangle\}$ does not include $(N \pm 1)$ states where in addition to adding or removing an electron, other electrons are excited from the ground state, i.e., terms involving more than one creation or annihilation operator. Thus we do not expect to be able to realistically describe quasiparticle weights, lifetimes, or satellite features.

Replacing in Eq. (4) the complete set of exact eigenfunctions by this incomplete set of approximate eigenfunctions [in Eq. (5)] and using the fact that the natural orbitals diagonalize γ , i.e., $\langle \Psi_0^N | a_{\lambda'}^{\dagger} a_{\zeta} | \Psi_0^N \rangle = \delta_{\lambda \zeta} n_{\zeta}$, we end up with the following approximation for the spectral function:

$$A_{\lambda\lambda'}(\omega) = 2\pi\delta_{\lambda\lambda'}[n_{\lambda}\delta(\omega - \epsilon_{\lambda}^{-}) + (1 - n_{\lambda})\delta(\omega + \epsilon_{\lambda}^{+})],$$
(6)

with $\epsilon_{\lambda}^{\pm} = E_0^N - E_{\lambda}^{N\pm 1}$. We note that in spite of being approximate, the spectral function in Eq. (6) satisfies the exact sum rule, $\frac{1}{2\pi} \int_{\infty}^{\infty} A_{\lambda\lambda'}(\omega) d\omega = 1$. Being, as function of ω , a single δ function for each fixed $\lambda = (i, \mathbf{k})$, the spectral function in Eq. (6) is reminiscent of a noninteracting meanfield type approximation. We emphasize that our approximation is by no means mean field because the $(N \pm 1)$ states in Eq. (5) are correlated and may even be strongly correlated if, for example, $|\Psi_0^N\rangle$ represents the ground state of a Mott insulator.

It is a formidable task to determine quasiparticle lifetimes within the Lehmann representation: one needs to determine the position of an infinite number of peaks corresponding to an infinite number of $N \pm 1$ eigenstates in Eq. (4). This coalescence of peaks is described by an envelope function the width of which is proportional to the inverse quasiparticle lifetime. In order to determine these lifetimes within the present formalism, one may use the more general form of the $N \pm 1$ states known from the extended Koopmans' theorem [18], $|\chi_i^{N-1}\rangle =$ $\sum_{\lambda} \alpha_{i\lambda}^{\pm} |\Phi_{\lambda}^{N\pm1}\rangle$, where $|\Phi_{\lambda}^{N\pm1}\rangle$ are given by Eq. (5). This yields for the diagonal of the spectral function, $A_{\lambda\lambda}(\omega) =$ $2\pi[n_{\lambda}W_{\lambda}^{-}(\omega) + (1 - n_{\lambda})W_{\lambda}^{+}(\omega)] \quad \text{with} \quad W_{\lambda}^{\pm}(\omega) = \sum_{j}|\alpha_{j\lambda}^{\pm}|^{2}\delta(\omega \pm [E_{0}^{N} - E_{j}^{N\pm 1}]), \text{ which has a finite width}$ possibly allowing us to determine the quasiparticle lifetimes. Such investigation will be left to the future. Within this article we wish to focus on the so-called density of states, which is obtained by taking the trace of the spectral function $(\sum_{\lambda} A_{\lambda\lambda})$:

$$DOS = 2\pi \sum_{\lambda} [n_{\lambda} \delta(\omega - \epsilon_{\lambda}^{-}) + (1 - n_{\lambda}) \delta(\omega + \epsilon_{\lambda}^{+})].$$
(7)

Now what remains is to calculate the excitation energies $\epsilon_{\lambda}^{\pm} = \epsilon_{i\mathbf{k}}^{\pm} = E_0^N - E_{i\mathbf{k}}^{N\pm 1}$, where $E_{i\mathbf{k}}^{N\pm 1}$ is the energy of the system with an electron, with specific momentum \mathbf{k} , added or removed. While in experiments $E_{i\mathbf{k}}^{N\pm 1}$ represents the total energy of a macroscopic block of material, in the theoretical description $E_{i\mathbf{k}}^{N\pm 1}$ is the total energy of a large but periodically repeated Born–von Karman (BvK) cell, where a constant charge background is added to keep the total (infinite) system charge neutral. This would require a prohibitively large number of total-energy calculations-equal

to the number of **k** points times the number of natural orbitals (typically ~2500). This is a formidable task and hence we make another simplification which is *not conceptual* in nature but rather a numerical trick similar to the Slater transition state procedure [19]; we first introduce the total ground-state energy, $E_{i\mathbf{k}}^{N\pm\eta}$, where a fractional number of particles, η , has been added or removed at a given $i\mathbf{k}$. We then assume that upon adding or removing charge to $i\mathbf{k}$ the only occupation number that changes significantly is the one that corresponds to the very same $i\mathbf{k}$ while all the other occupation numbers as well as the natural orbitals remain unchanged. With this simplification, following Slater [19], the ϵ^{\pm} , can be approximated as

$$\epsilon_{i\mathbf{k}}^{\pm} = \frac{\partial E_{i\mathbf{k}}^{N\pm\eta}}{\partial\eta} \Big|_{\eta=1/2} = \frac{\partial E[\{\phi\},\{n\}]}{\partial n_{i\mathbf{k}}} \Big|_{n_{i\mathbf{k}=1/2}}.$$
 (8)

This simplification can be easily numerically validated by plotting the total energy as a function of n_{ik} ; we find a nearly linear behavior for all the materials studied here. This implies that the Slater-type evaluation of the total-energy difference in Eq. (8) is rather accurate.

Following the above procedure the spectral density for the strongly correlated Mott insulators NiO, CoO, FeO, and MnO is calculated using the full-potential linearized augmented plane wave (FP-LAPW) code Elk [20], with practical details of the calculations following the scheme described in Ref. [11].

It is immediately apparent from Fig. 1 that RDMFT captures the essence of Mott-Hubbard physics: all the TMOs considered are insulating in nature. This fact was already noticed in the previous work [11] where the presence of a gap without any spin order was deduced via a very different technique, namely the discontinuity in the



FIG. 1 (color online). Density of states for the TMOs in the presence of antiferromagnetic order. Site and angular momentum projected spectral density are also presented for transition metal e_g and t_{2g} states and oxygen-*p* states. In addition, XPS and BIS spectra (shifted up for clarity) are presented for comparison [22–26]. Again, $\alpha = 0.656$ for all materials.

chemical potential as a function of the particle number. A closer examination of the spectra for NiO, CoO, and MnO reveals a good agreement between the RDMFT peaks and the corresponding x-ray photoemission spectroscopy (XPS) and bremsstrahlung isochromat spectroscopy (BIS) data. In fact, not only the peak positions, but also their relative weights are well reproduced. For the case of FeO, it must be recalled that Fe segregation, unavoidable in this compound, precludes the experimental realization of pure FeO samples. For this reason the only existing experimental data are rather old and the presumably substantially contaminated and broadened data present no distinct features that may be used for a clear comparison.

The actual values of the insulating gaps that may be extracted from Fig. 1 are 4.5(4.3) eV, 2.6(2.8) eV, 3.2 (3.6) eV, 3.5 eV for NiO, CoO, MnO, and FeO, respectively, with the corresponding experimental gap given in parentheses. The value of the local moments we find to be $1.36(1.9)\mu_B$, $2.7(3.3)\mu_B$, $3.35(3.62)\mu_B$, and $3.38(4.7)\mu_B$ for NiO, CoO, FeO, and MnO, respectively, again with the experimental values in parentheses. There are two reasons for the smaller values of the magnetic moment within RDMFT compared to experiment. First, the calculations are performed with the FP-LAPW method in which space is divided into spheres around the atoms, the so-called muffin tins, and the interstitial region. In the case of fully noncollinear magnetic calculations the magnetic moment per site is calculated by integrating the magnetization vector field inside the muffin tin. This means a small part of the moment is lost in the interstitial. Second, the power functional induces a slight noncollinearity in the magnetization leading to yet more loss in the z-projected moment.

In Fig. 1 we also present the site and angular momentum projected spectral density for the TMOs considered in this work. The electronic gap, as expected, always occurs between lower and upper Hubbard bands dominated by transition metal d states. However, while for NiO one finds a significant component of oxygen-p states in the lower Hubbard band, for the other TMOs this hybridization between oxygen-p and TM-d states reduces, and is almost absent, in the case of MnO, indicating that for this material the insulating state is driven mostly by Mott-Hubbard correlations. As a validation of our method for the calculation of spectra we may compare these features of the projected spectral density, and in particular the ordering in energy of the t_{2g} and e_g states, with well-established *ab initio* many-body techniques such as DMFT and the GW method [1,2]. In all cases we find a very good agreement, signaling that the method we present here yields an accurate description of the detailed features of the spectral density.

Theoretical methods used to study TMOs at zero temperature and pressure agree with each other as far as the spectral density of TMOs is concerned [1,2,10,21]; however, the actual values of the gap and the moment differ depending upon the details of the calculations. However, this agreement between various methods ends at ambient conditions, in that when attempting to study pressure induced insulator-metal phase transitions (IMT) the results vary wildly depending on the method. In MnO, experimental data point toward a first order IMT which is accompanied by a reduced volume ($v/v_0 = 0.68$ to 0.63) and moment collapse (5 to $1\mu_B$) [3]. DFT with local density approximation/generalized gradient approximation (LDA/ GGA)-like functionals captures this collapse of volume and magnetic moment and shows that the physics behind this phase transition is the simple widening of the Mn dstates due to increased itinerancy of the electrons at the reduced volume [9,10]. On the other hand, DMFT shows that the phase transition occurs as a result of the increased crystal field splitting with the width of the Mn d states unchanged [1]. The picture obtained with correlated band theories is even more complicated: LDA + U shows a moment collapse but no IMT [21], hybrid functionals yield a phase transition to a semimetalic state [21], and finally the self-interaction corrected DFT shows a transition to a metallic state with unusually large *d*-band width [21].

Given this wide spread of results, in the present work we study insulator to metal phase transitions using RDMFT to shed light on this controversy. RDMFT is an ideal method for doing so as it is an *ab initio* theory for treating strongly correlated systems not requiring any adjustable system dependent parameters. The results, obtained using RDMFT, for the magnetic moment in MnO under applied pressure are shown in Fig. 2. It is clear that the magnetic moment collapses from $3.6\mu_B$ at optimal volume to $0.54 \mu_B$ at a reduced volume. Further reduction of the volume does not change the moment. Within RDMFT we find a volume collapse of 11%, which is higher than the experimental value of 6.6%. In order to investigate the prime reason behind this moment collapse we plot in Fig. 2 the number of electrons in the Mn d states as a function of volume. One notices a redistribution of



FIG. 2. Left panel: on-site magnetic moment (in μ_B) for MnO as a function of reduced volume (ν/ν_0). Right panel: Mn *d*-band occupancy resolved into e_g and t_{2g} components.

electrons among the symmetry projected t_{2g} and e_g states. At the reduced volume of 0.8 the number of e_g electrons (n_{eg}) starts to reduce, finally leveling off at $v/v_0 = 0.711$. This is accompanied by an increase in the t_{2g} state charge (n_{t2g}) . This picture is fully concomitant with the previous results obtained using DMFT [1]. In order to look at the detailed behavior of these t_{2g} and e_g states we have also plotted them as a function of volume in Fig. 3. The e_g states move above the chemical potential as the volume is reduced, while the t_{2g} states move below, finally ending up with a totally different spectral density in the metallic phase as compared to the Mott insulating phase. As a result of this rearrangement the crystal field splitting between the t_{2g} and e_g states increases. Despite RDMFT being a totally different approach as compared to DMFT, the symmetry projected spectral density as a function of volume looks very similar for the two methods (see Fig. 3 of Ref. [1]). However, there exists a striking difference between the two in that within RDMFT the Mn d states widen at reduced volume as compared to those at $v/v_0 = 1$. This increase in the bandwidth has its origin in a small shift in the spectral weight of the lower Hubbard band to lower energy. A close inspection of Fig. 3 shows that the change in crystal field splitting is very large ($\sim 4.4 \text{ eV}$) and certainly has a prominent role in the IMT in MnO. Band widening is a coexisting phenomenon that has a very small influence on the IMT. This band widening on the other hand drives the IMT if LDA/GGA functionals with DFT are used.

To conclude, we have presented a method to calculate photoelectron spectra within the framework of RDMFT. We have shown that the spectral information obtained in this way gives a detailed account of the strongly correlated nature of the TMOs, including the subtle interplay between



FIG. 3 (color online). t_{2g} (right panel) and e_g (left panel) projected spectral density for MnO as a function of energy (in eV) and reduced volume. Results are obtained using RDMFT. Spectral density for different reduced volumes are shifted vertically for clarity.

Mott-Hubbard correlation and charge-transfer character in these materials. We validate this method not only by showing a good agreement of gross spectral features with experiments, but also by a detailed comparison of the angular momentum projected spectral density with that of well-established many-body techniques, in all cases finding a very good agreement. We have further elucidated the physics behind the insulator to metal phase transition in MnO using RDMFT. For MnO the pressure induced phase transition is caused by the increase in crystal field splitting, which in turn is the result of a redistribution of charge among the states with t_{2g} and e_g symmetry. The widening of the transition metal *d* band is seen as a coexisting phenomenon but is certainly not the reason for the metallization.

*sharma@mpi-halle.mpg.de

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